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New type of transients in the era of all-sky surveys





OPTICON RadioNet Pilot





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Era of all-sky surveys

- Current online: surveys ASAS-SN, ATLAS, Gaia, GOTO, Pan-STARRS, OGLE and ZTF
- Current 'delay'-online: VMC, VVV and WISE
- Very soon: Rubin/LSST and Roman/WFIRST



ULIRG F01004-2237



F01004-2237



- HST/STIS spectrum (2000) fitted with WN7-9 + WC7-8 + blueshifted HeII λ4686 (NLR)
- WHT/ISIS spectrum (2015) required additional HeII λ4686 component with FWHM~6200 km/s.
- CL-AGN was rule out because of presence of HeII/NIII component

OGLE17aaj

z=0.116 $M_v = -18.3$ $M_{\rm BH} \sim 2.5 \ {\rm x} \ 10^7 \ M_{\rm Sun}$











Gromadzki et al. 2019

AT 2017bgt (ASASSN-17cv)



Trakhtenbrot et al. 2019

z = 0.064 $M_{\rm v} \sim -21.3$

AT 2017bgt





Trakhtenbrot et al. 2019

Bowen fluorescence





Netzer et al. 1985

AT 2017bgt: evolution



Trakhtenbrot et al. 2019

AT 2017bgt: from X-ray to NIR



Trakhtenbrot et al. 2019

Gaia2019axp (AT 2019brs)



z = 0.374 $M_{\rm G} \sim -22.3$

Ihanec et al. in prep





Impact on TDEs Science

- Leloudas et al. 2019 found Bowen features in spectra of AT 2019dyb and also proved that previous TDEs also showed them
- Blagorodnova et al. 2019 found
 Bowen features in spectra of iPTFaf
- Van Velzen et al. 2021 proposed three classes of TDEs: (i) TDE-H,
 (ii) TDE-Bowen and (iii) TDE-He



Few take-home remarks

- Still no satisfactory scenario proposed
- Masses of BHs look similar ($10^6 10^7 M_{Sun}$);
- Wide range of released energy $(L_{\text{peak}} \sim 10^{42} 10^{44} \text{ erg s}^{-1});$
- Different spectroscopic and photometric evolution
- Few new similar transient were published and also discovered in the meantime, eg. ASAS-SN18jd, AT 2019avd
- Two of them show second re-brightening: F01004-2236 and AT 2019avd



